

1) # modes with wavenumber less than k is

$$N(k) = \left(\frac{L}{2\pi}\right)^3 \frac{4}{3} \pi k^3 \quad \begin{matrix} 2 \\ \uparrow \\ 2 \text{ electron spins} \end{matrix}$$

$$\text{But } E = \frac{\hbar^2 k^2}{2m} \Rightarrow N(E) = \left(\frac{L}{2\pi}\right)^3 \frac{4}{3} \pi \left(\frac{2mE}{\hbar^2}\right)^{3/2} 2$$

\therefore energy density of electron states is:

$$D(E) = \frac{dN}{dE}$$

$$= \left(\frac{L}{2\pi}\right)^3 4\pi \left(\frac{2m}{\hbar^2}\right)^{3/2} E^{1/2}$$

electrons $N \equiv \int_0^{\epsilon_F} D(E) dE$ This is definition of Fermi energy.

$$N = \int_0^{\epsilon_F} \left(\frac{L}{2\pi}\right)^3 4\pi \left(\frac{2m}{\hbar^2}\right)^{3/2} E^{1/2} dE$$

$$= \frac{V}{2\pi^2} \left(\frac{3m}{\hbar^2}\right)^{3/2} \frac{2}{3} \epsilon_F^{3/2}$$

energy at $T=0$ $U_0 \equiv \int_0^{\epsilon_F} E D(E) dE$.

$$= \left(\frac{L}{2\pi}\right)^3 4\pi \left(\frac{2m}{\hbar^2}\right)^{3/2} \int_0^{\epsilon_F} E^{3/2} dE$$

$\therefore U_0 = \frac{3}{5} N \epsilon_F$ using expression for N

$$4a) M_s = 2 \times 10^{33} \text{ gm.}$$

Assuming sun is all H; # electrons in sun = # protons.

$$\begin{aligned} \therefore N &= \frac{M_s}{m_{\text{prot.}}} \\ &= \frac{2 \times 10^{33}}{1.67 \times 10^{-24}} \\ &= 1.2 \times 10^{57} \text{ electrons} \end{aligned}$$

White Dwarf; $r = 2 \times 10^9 \text{ cm.}$

$$\begin{aligned} E_F &= \left(\frac{3\pi^2 N}{V} \right)^{2/3} \frac{\hbar^2}{3m} \\ &= \left(\frac{3\pi^2 \cdot 1.2 \times 10^{57}}{\frac{4\pi}{3} (2 \times 10^9)^3 \text{ cm}^3} \right)^{2/3} \frac{(1 \times 10^{-27} \text{ erg sec})^2}{3 \times 9.11 \times 10^{-28} \text{ gm.}} \end{aligned}$$

$$= 9.8 \times 10^{-9} \text{ erg.}$$

$$\therefore E_F = 6100 \text{ eV.}$$

b) # electron states having wavevector less than k is

$$N(k) = \left(\frac{L}{2\pi} \right)^3 \frac{4\pi k^3}{3} \cdot 2$$

$$\text{Using } E = \hbar k c \Rightarrow N(E) = \frac{V}{3\pi^2} \frac{E^3}{(\hbar c)^3}$$

$$\therefore \text{density of states } D(E) = \frac{V}{\pi^2} \frac{E^2}{(\hbar c)^3}$$

$$N = \int_0^{\epsilon_F} g(\epsilon) d\epsilon \quad \text{definition of Fermi level}$$

$$= \frac{V}{3\pi^2} \frac{\epsilon_F^3}{(\hbar c)^3}$$

$$\therefore \epsilon_F = \left(\frac{3\pi^2 N}{V} \right)^{1/3} \hbar c$$

c) If $r = 10^6 \text{ cm}$
 $N = 1.2 \times 10^{57}$ electrons

$$\Rightarrow \epsilon_F = \left(\frac{3\pi^2 \times 1.2 \times 10^{57}}{\frac{4\pi}{3} \times 10^{18} \text{ cm}^3} \right)^{1/3} \times 1 \times 10^{-27} \text{ erg sec} \times 3 \times 10^{10} \frac{\text{cm}}{\text{sec}}$$

$$= 6.5 \times 10^{-4} \text{ erg}$$

$$= 4 \times 10^8 \text{ eV}$$

5) Fermi Energy for liquid He^3 is:

$$\epsilon_F = \left(\frac{3\pi^2 N}{V} \right)^{2/3} \frac{\hbar^2}{2m} \quad \text{where } \frac{N}{V} = \frac{Nm}{mV} \quad m = 3 m_{\text{prot}}$$

$$= \frac{p}{m}$$

$$\epsilon_F = \left(\frac{3\pi^2 \times 0.081 \text{ gm/cm}^3}{3 \times 1.67 \times 10^{-24} \text{ gm}} \right)^{2/3} \frac{(1 \times 10^{-27} \text{ erg sec})^2}{2 \times 2 \times 1.67 \times 10^{-24} \text{ gm}}$$

$$= 0.7 \times 10^{-15} \text{ erg}$$

$$= 4.2 \times 10^{-4} \text{ eV}$$

$$T_F = \frac{\epsilon_F}{k} \approx 5 \text{ K.}$$

$$b) \quad m \frac{dv}{dt} + \frac{mv}{\tau} = -eE.$$

$$\text{let } E = E e^{-i\omega t}$$

$$v = v e^{-i\omega t}$$

$$\therefore -im\omega v + \frac{mv}{\tau} = -eE.$$

$$mv \left(\frac{1}{\tau} - i\omega \right) = -eE.$$

$$v = -\frac{eE}{m} \frac{1}{\frac{1}{\tau} - i\omega} \\ = -\frac{e\tau}{m} \frac{1}{1 - i\omega\tau} E.$$

$$\text{Current Density } J = -nev.$$

$$= \frac{ne^2\tau}{m} \frac{1}{1 - i\omega\tau} E.$$

$$\therefore \text{conductivity } \sigma = \frac{ne^2\tau}{m} \frac{1 + i\omega\tau}{1 + \omega^2\tau^2}.$$

$$\therefore \sigma(\omega) = \sigma(0) \frac{1 + i\omega\tau}{1 + \omega^2\tau^2}$$

9) Equation of motion $\vec{B} = B\hat{z}$ is:

$$m \left(\frac{d}{dt} + \frac{1}{\tau} \right) v_x = -e \left(E_x + \frac{B}{c} v_y \right)$$

$$m \left(\frac{d}{dt} + \frac{1}{\tau} \right) v_y = -e \left(E_y - \frac{B}{c} v_x \right)$$

$$m \left(\frac{d}{dt} + \frac{1}{\tau} \right) v_z = -e E_z.$$

In the static case, all time derivatives are zero.

$$\therefore \frac{m}{\tau} v_x = -e \left(E_x + \frac{B}{c} v_y \right) \quad (1)$$

$$\frac{m}{\tau} v_y = -e \left(E_y - \frac{B}{c} v_x \right) \quad (2)$$

$$\frac{m}{\tau} v_z = -e E_z \quad (3)$$

Subst. v_y from (2) into (1), we get:

$$\frac{m}{\tau} v_x = -e E_x - \frac{eB}{c} \frac{\tau}{m} (-e) \left(E_y - \frac{B}{c} v_x \right)$$

$$v_x = -\frac{e\tau}{m} E_x - \omega_c \frac{\tau^2}{m} (-e) \left(E_y - \frac{B}{c} v_x \right) \quad \omega_c \equiv \frac{eB}{mc}$$

$$= -\frac{e\tau}{m} E_x + \frac{e\tau^2}{m} \omega_c E_y - \tau^2 \omega_c^2 v_x.$$

$$v_x (1 + \omega_c^2 \tau^2) = -\frac{e\tau}{m} \left(E_x - \omega_c \tau E_y \right)$$

$$v_x = \frac{-e\tau/m}{1 + \omega_c^2 \tau^2} \left(E_x - \omega_c \tau E_y \right)$$

$$J_x = -ne v_x$$

$$= \frac{ne^2 \tau / m}{1 + \omega_c^2 \tau^2} (E_x - \omega_c \tau E_y)$$

$$J_x = \frac{\sigma(0)}{1 + \omega_c^2 \tau^2} (E_x - \omega_c \tau E_y)$$

Similarly one solves for $J_y + J_z$.

$$\Rightarrow \begin{pmatrix} J_x \\ J_y \\ J_z \end{pmatrix} = \frac{\sigma(0)}{1 + \omega_c^2 \tau^2} \begin{pmatrix} 1 & -\omega_c \tau & 0 \\ \omega_c \tau & 1 & 0 \\ 0 & 0 & 1 + \omega_c^2 \tau^2 \end{pmatrix} \begin{pmatrix} E_x \\ E_y \\ E_z \end{pmatrix}$$